

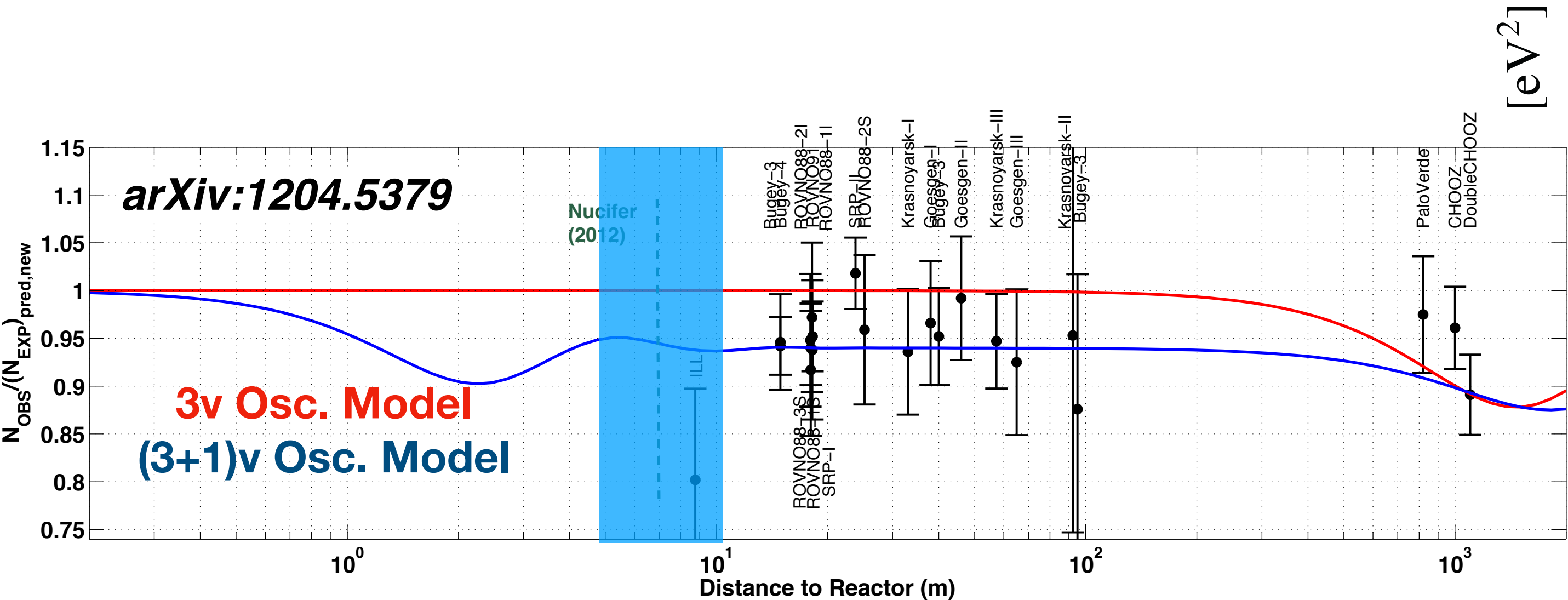
# **PROSPECT's latest results for Sterile Neutrino Oscillation search**

**Jose Palomino**  
**On behalf of Prospect collaboration**

**DNP 2020**

**October 30th, 2020**

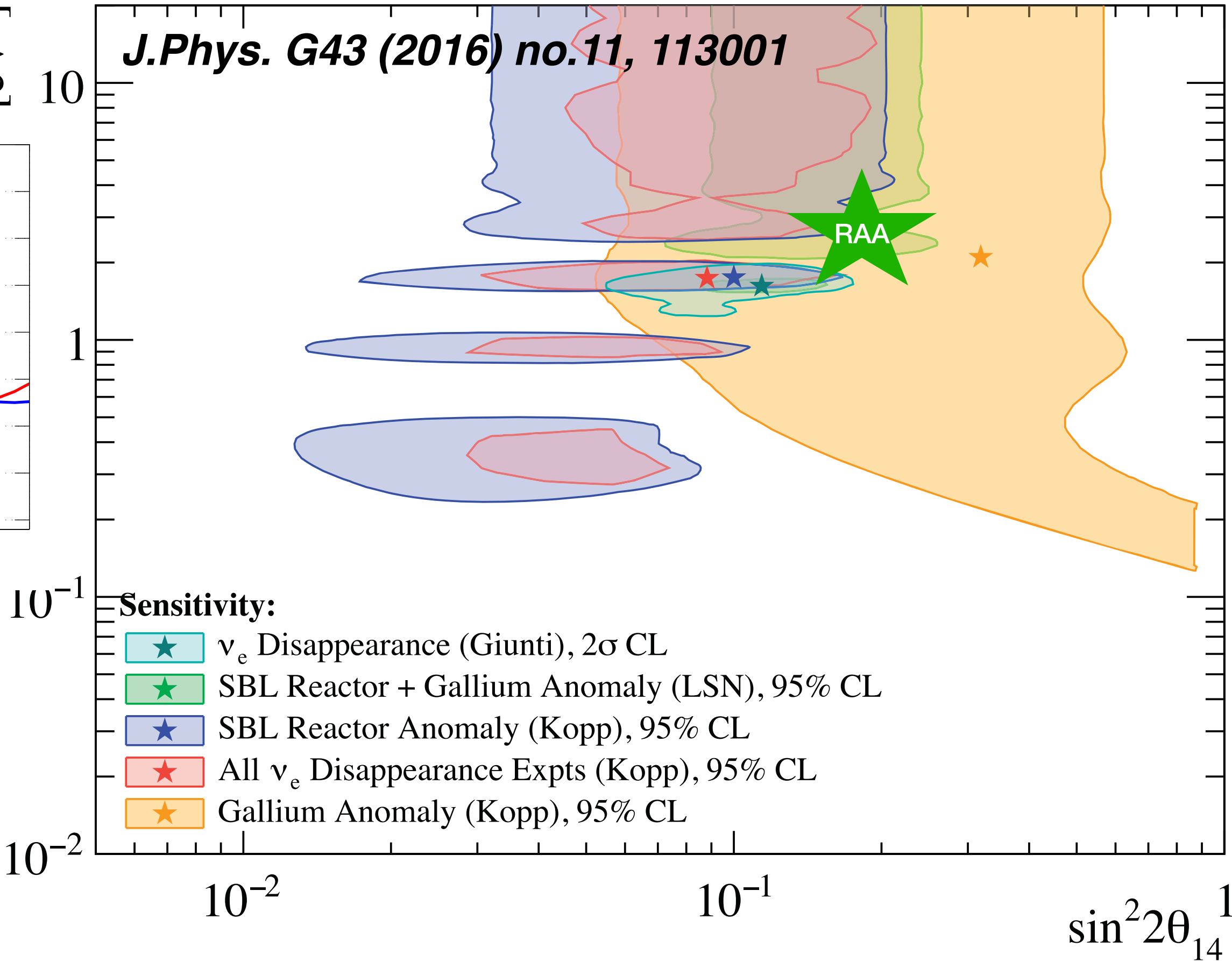
# Motivation: Reactor Antineutrino Anomaly (RAA)



- World average observed flux shows 6% deficiency with respect to theoretical predictions.
- The prediction models are based on Huber+Mueller and by 3-flavor neutrino oscillations at the distance of each experiment.

## Where this global deficit is coming from?

- Reactor model predictions are not good enough
- Sterile Neutrinos:
  - high frequency oscillations (~meter baselines).
  - eV-scale mass splitting.



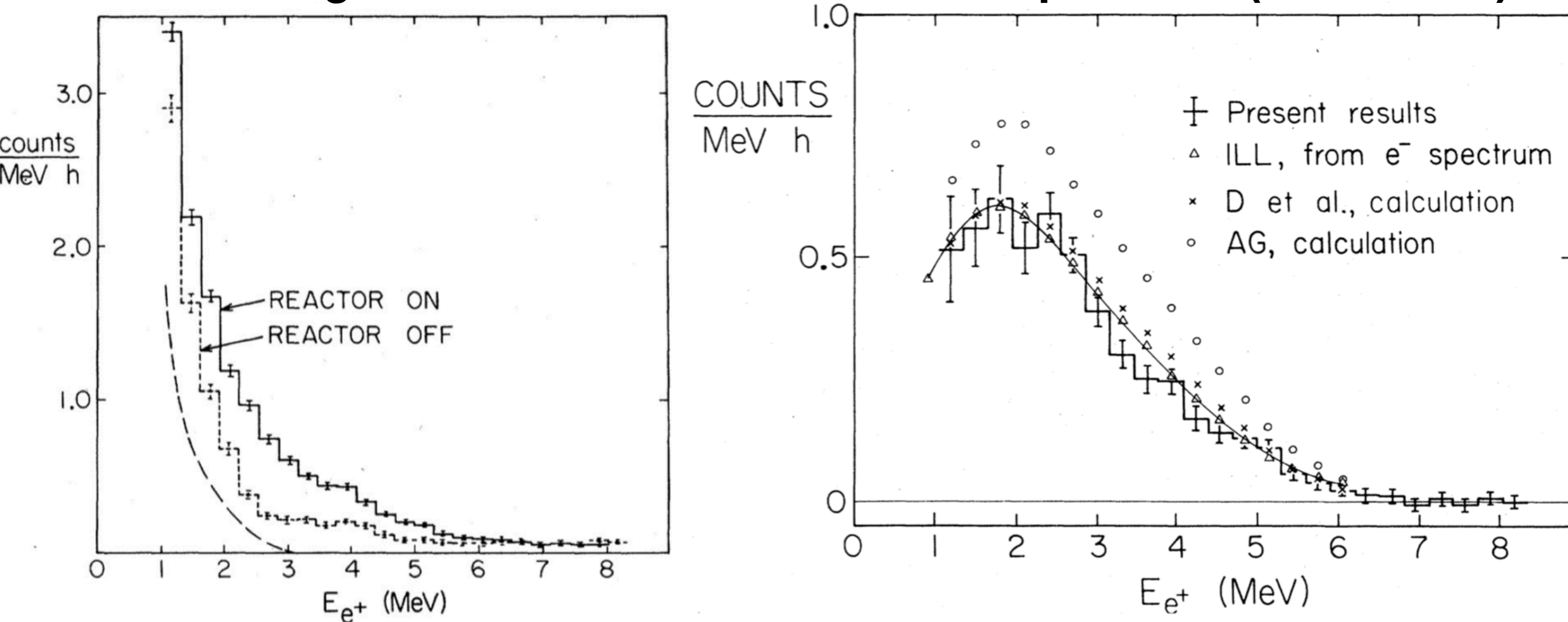
**RAA best-fit point at  $\sin^2 2\theta = 0.165$ ,  $\Delta m^2 = 2.39$**



# Physics Goals

There are not precise measurements at very short baseline.

Existing measurement from 1981 ILL experiment (~5k events).

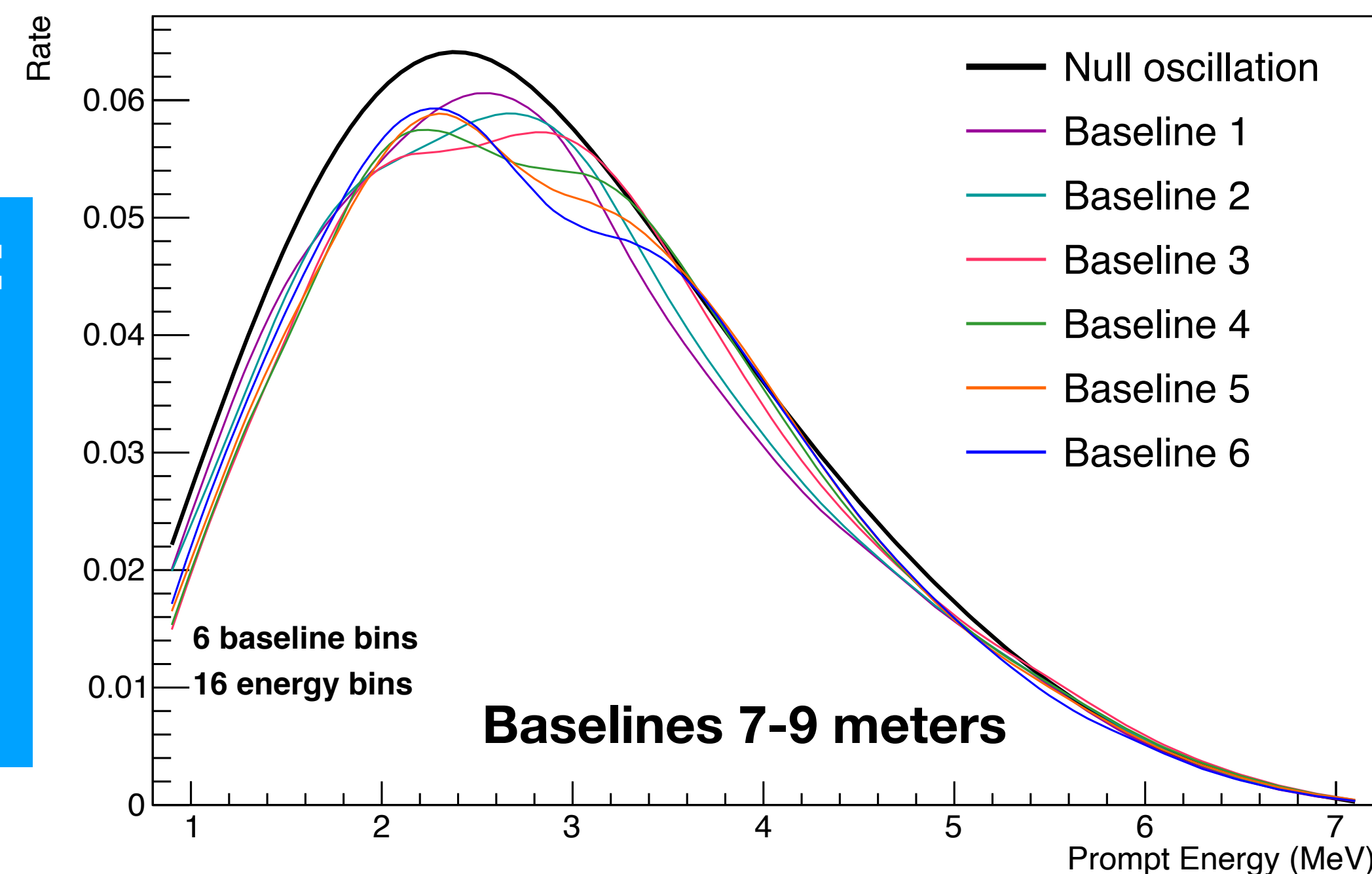


## Measurement of $^{235}\text{U}$ antineutrino spectrum:

- High energy resolution .
- High statistics.
- Have high enriched uranium cores:  $^{235}\text{U}$  only.

## Search for short-baseline sterile neutrinos:

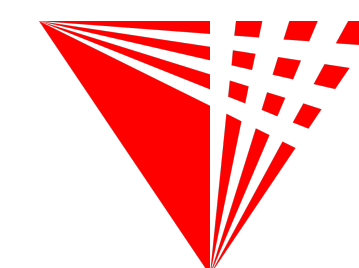
- Few meters baseline variation affects the predicted spectrum assuming sterile oscillations.
- Compact research reactor is necessary to prevent washing out oscillation.
- Reactor-model independent search for oscillations throughout the detector .







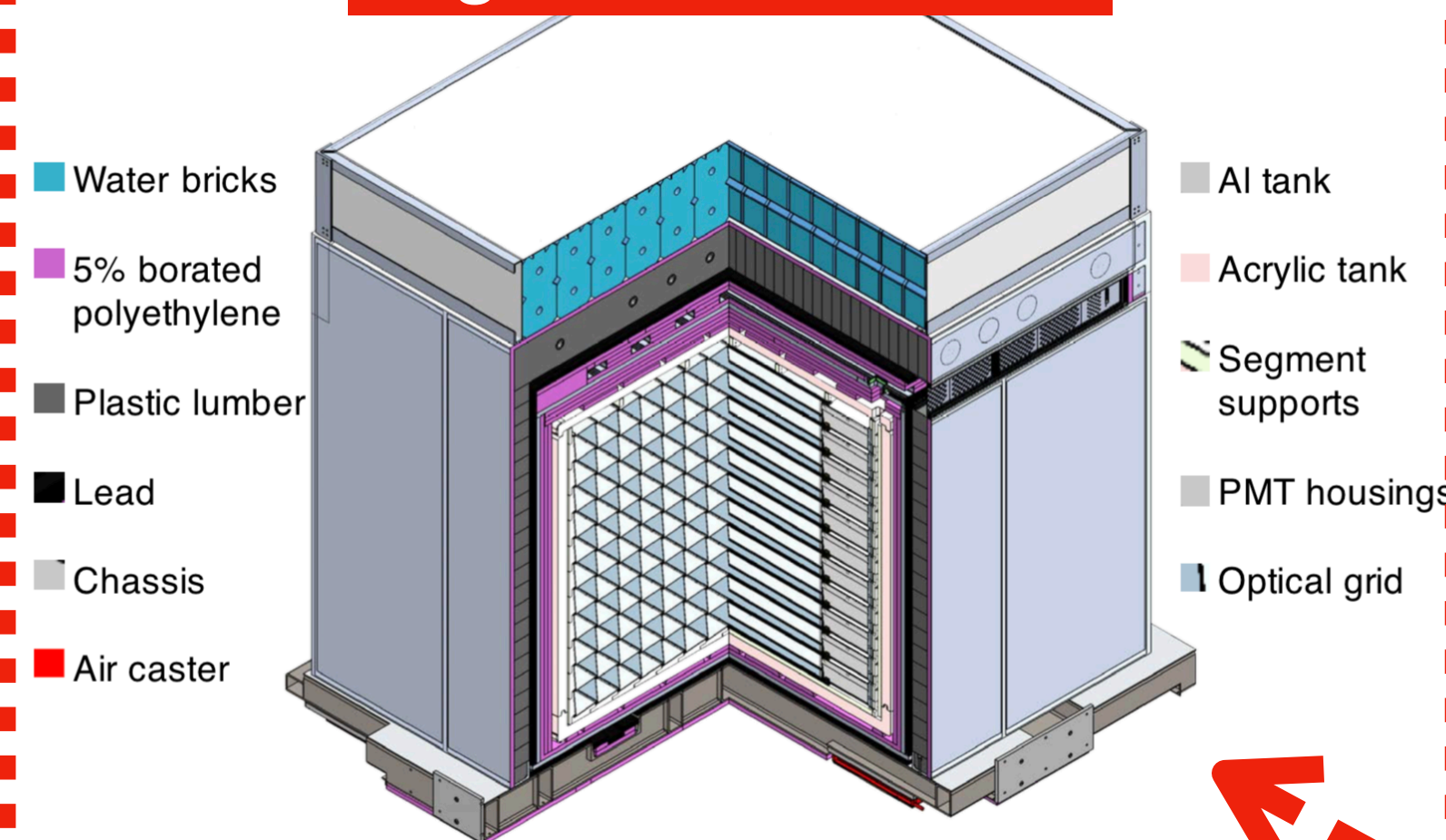
\*PROSPECT April 2020 Collaboration Meeting Photo



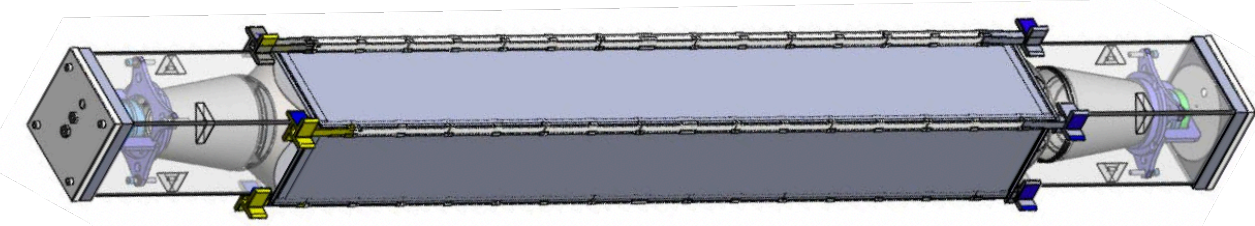


# Detector Design

## Segmented Detector



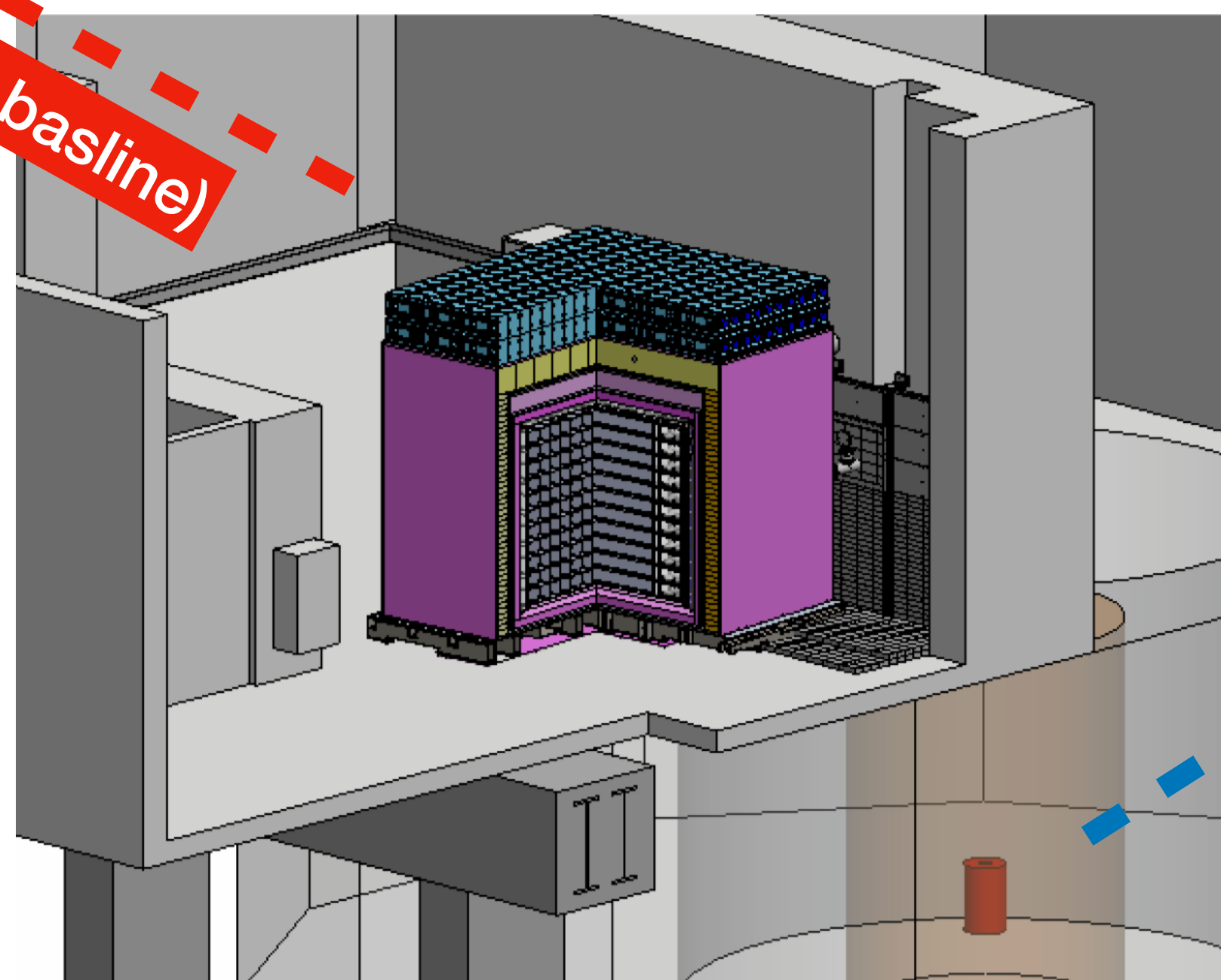
- ~3,000 L 6Li-loaded fiducial volume.
- 11 x 14 array of optically separated segments.
- Double ended PMT readout, with light concentrators.
- Good light collection and energy response  $\sim 4.5\text{-}5\%\sqrt{E}$  energy resolution.
- Full X,Y,Z event reconstruction.



## HIGH FLUX ISOTOPE REACTOR AT OAK RIDGE NATIONAL LABORATORY

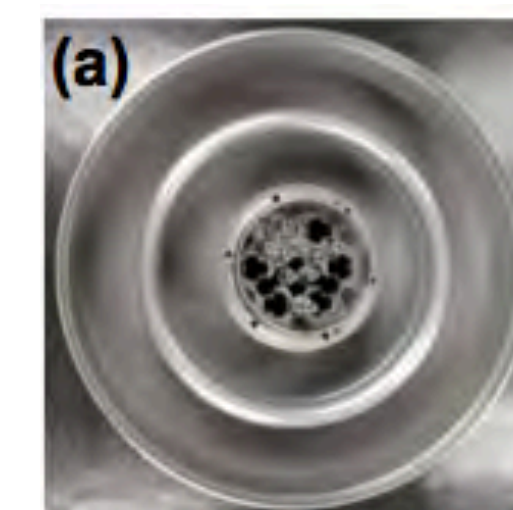


7m (short baseline)

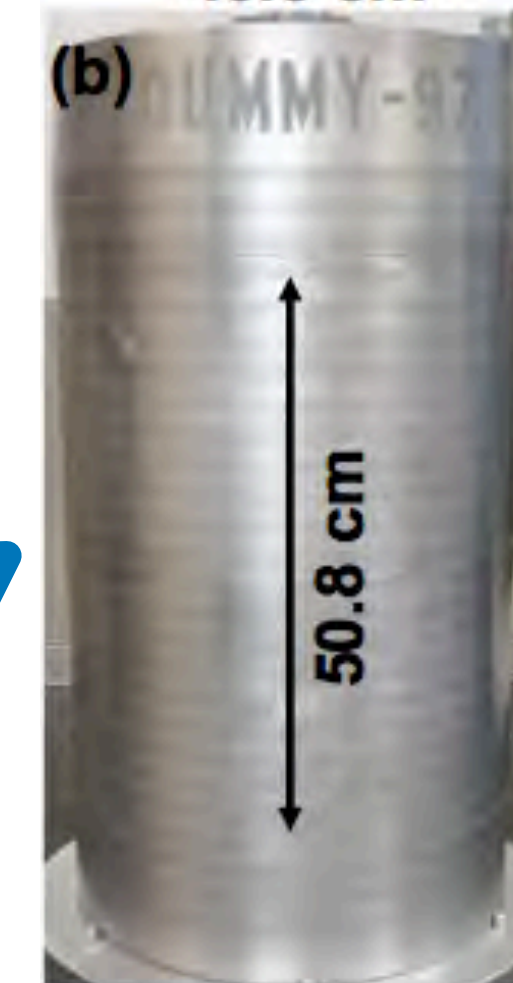


Reactor Core highly-enriched (HEU):  
>99% of  $\nu_e$  flux from  $^{235}\text{U}$  fission:

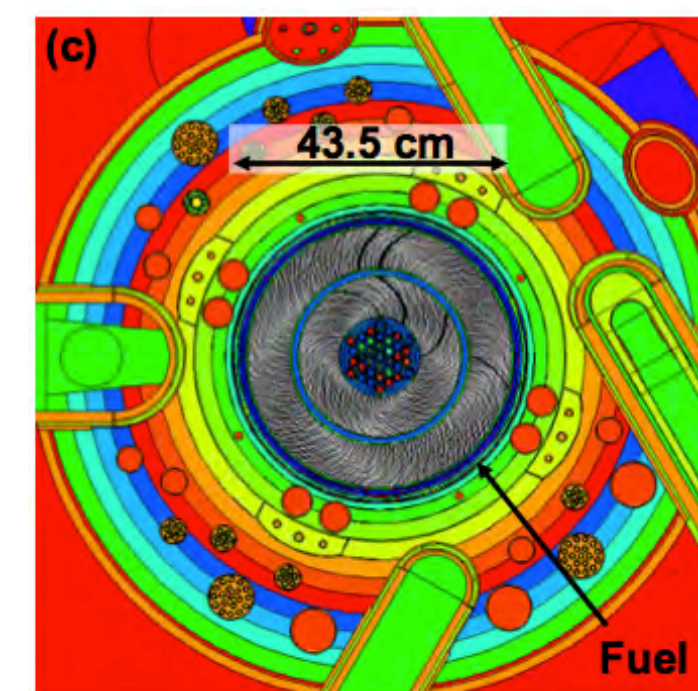
- Power: 85 MW
- Core shape: cylindrical
- Size:  $h=0.5\text{m}$   $d=0.4\text{m}$
- Duty-cycle: 24 days cycle



43.5 cm



50.8 cm

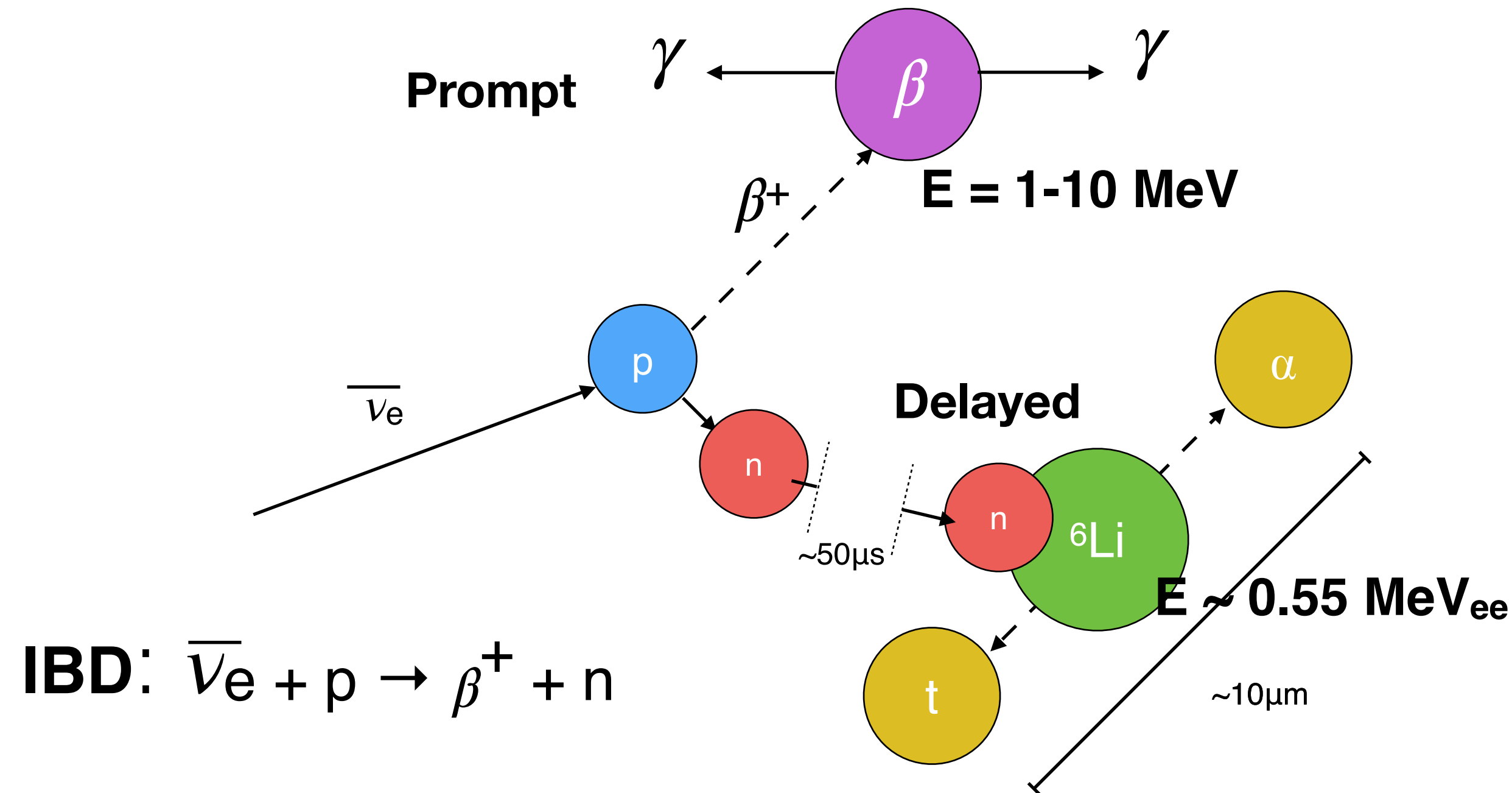


Fuel

Reactor is smaller than conventional power reactors



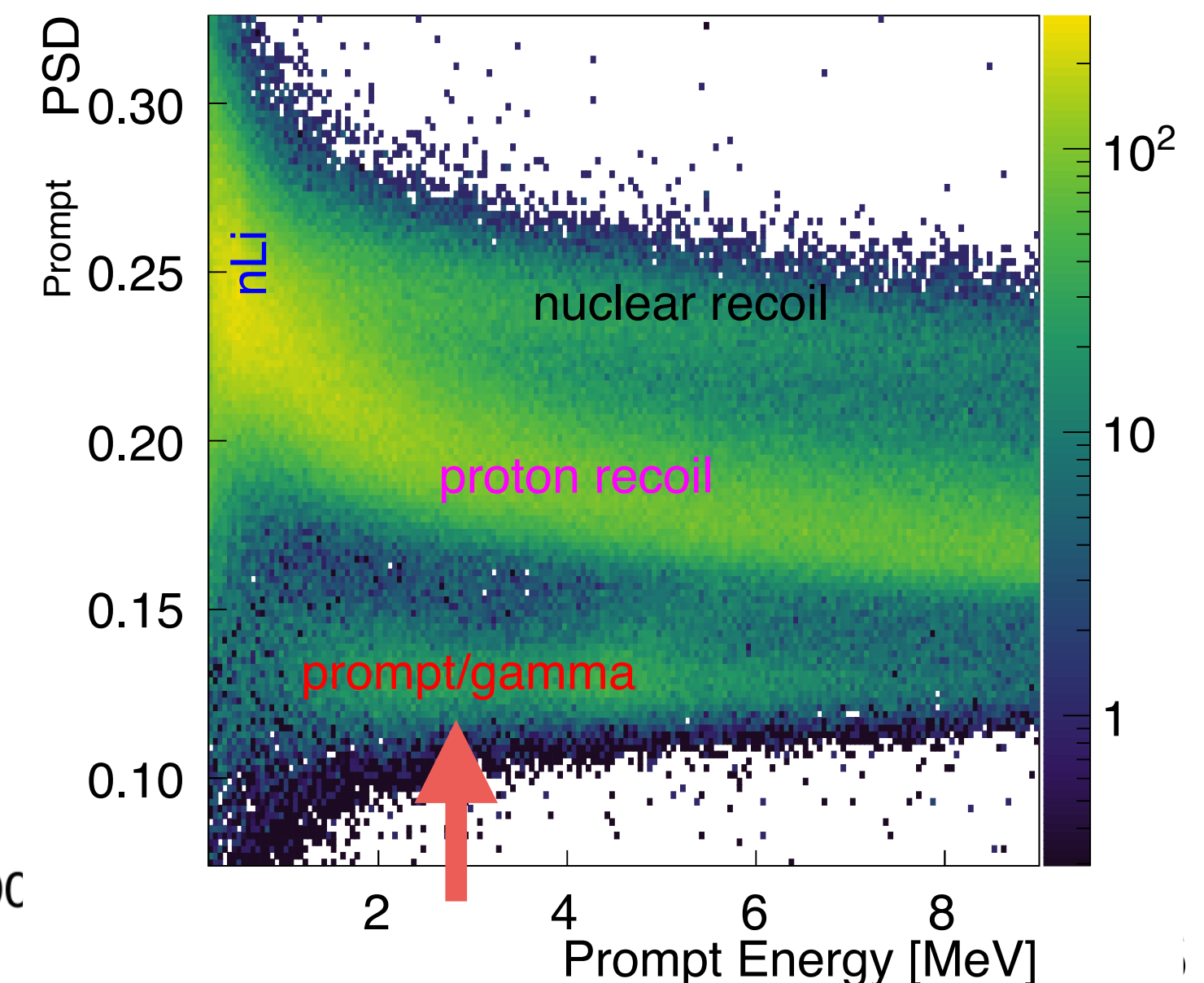
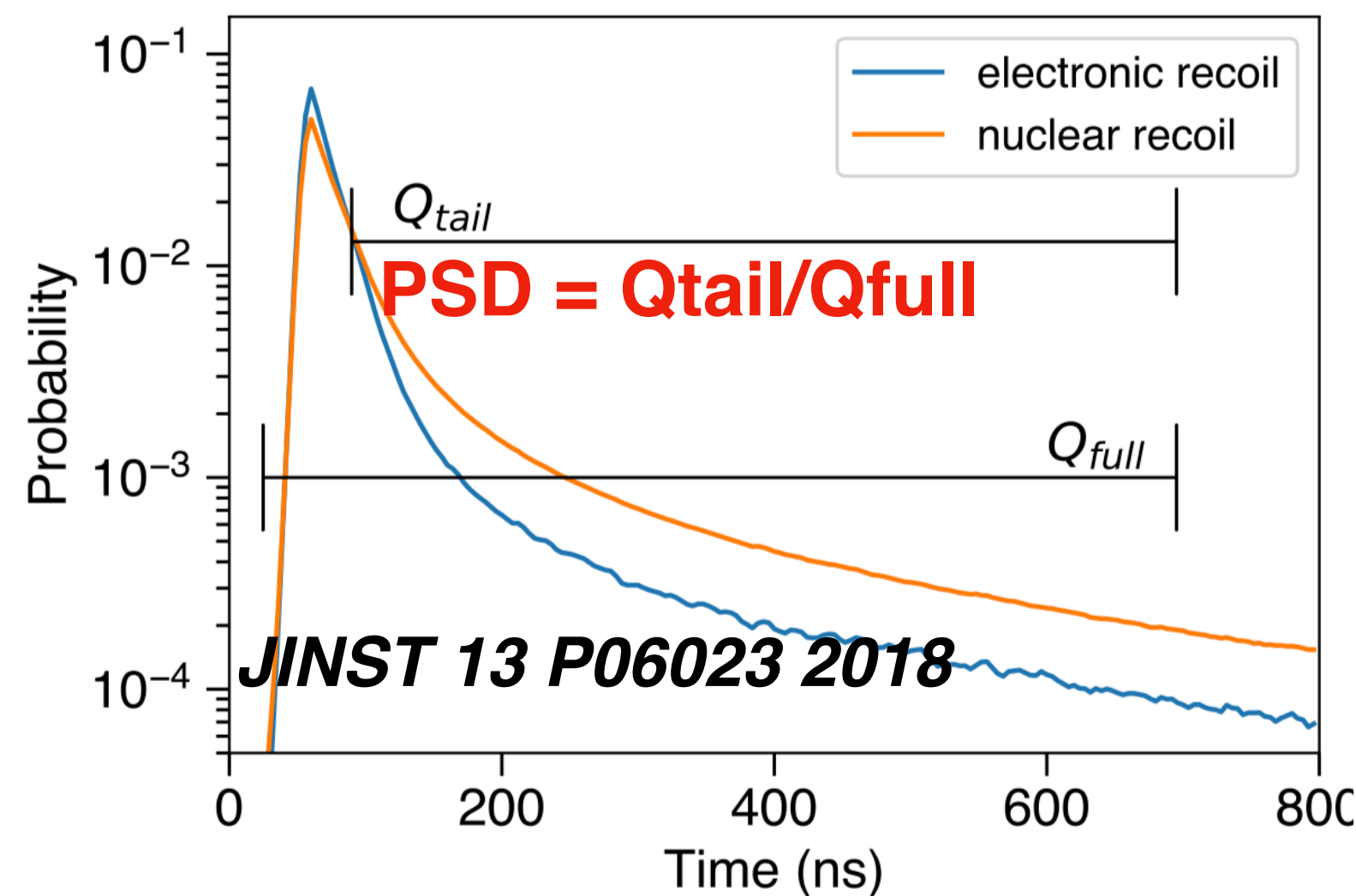
# IBD detection with ${}^6\text{LiLS}$



- 1-10 MeV  $\beta^+$ -like prompt signal (ionization and annihilation of positron).
- Followed by  $\sim 50 \mu\text{s}$  delayed neutron ( $\sim 0.55 \text{ MeV}$ ) capture on  ${}^6\text{Li}$ .
- ${}^6\text{LiLS}$  ideal for neutron tag in compact detector as decay is highly localized in space within a segment.

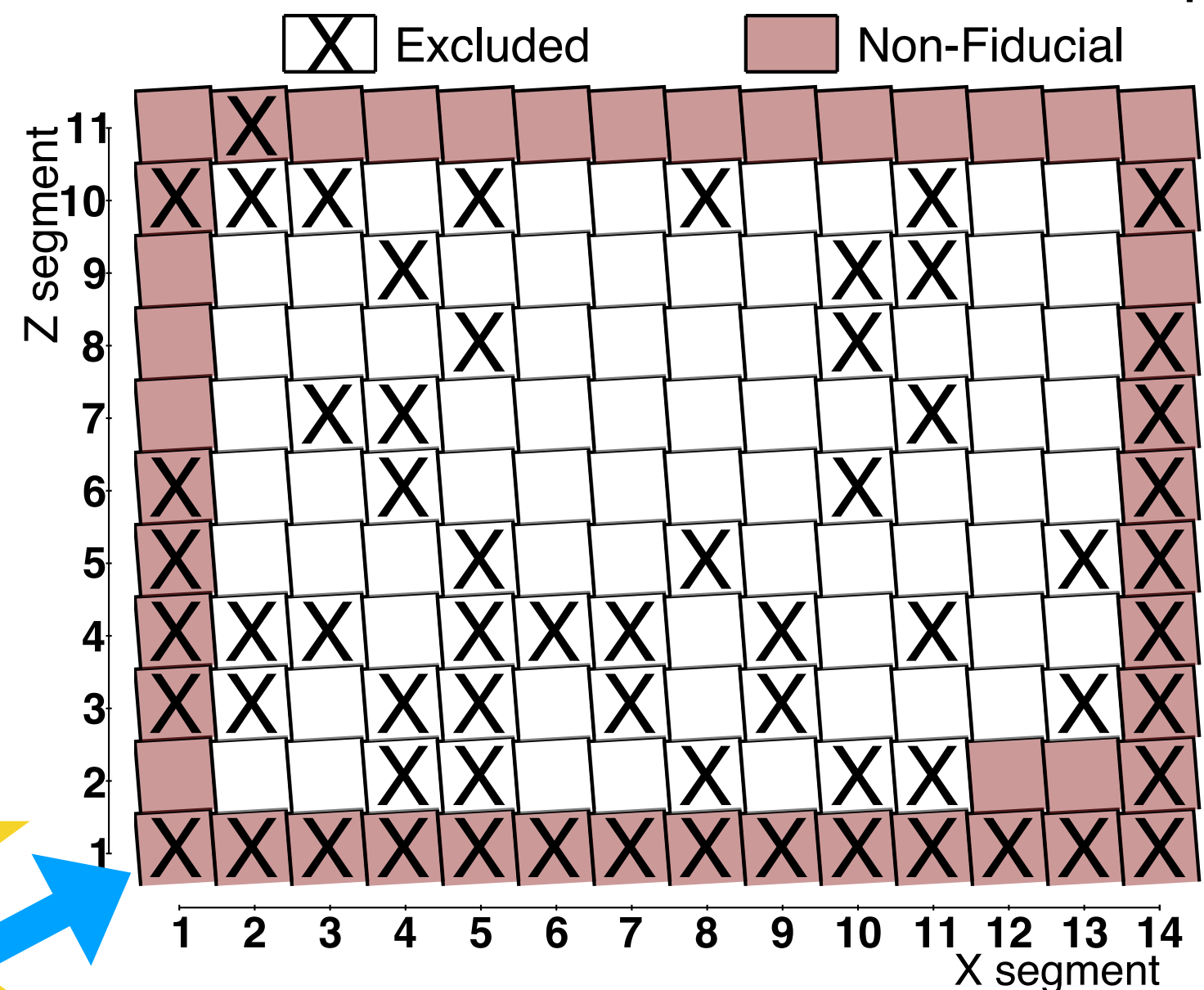
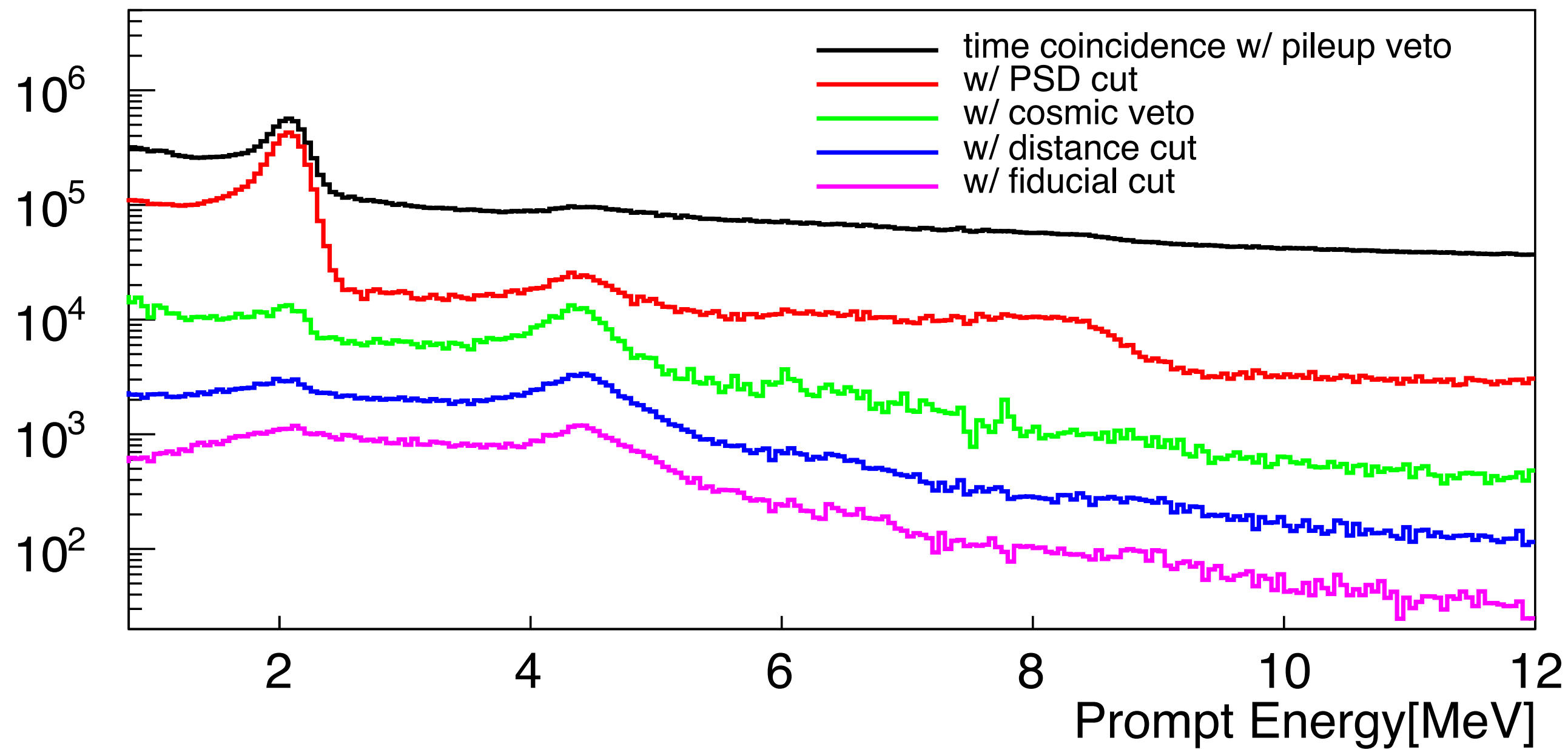
The Pulse Shape Discrimination (**PSD**) of scintillator works as particle identification.

- it can distinguishes gamma interactions, neutron capture and nuclear recoils.
- Essential to remove cosmogenic neutrons background.





# IBD Selection

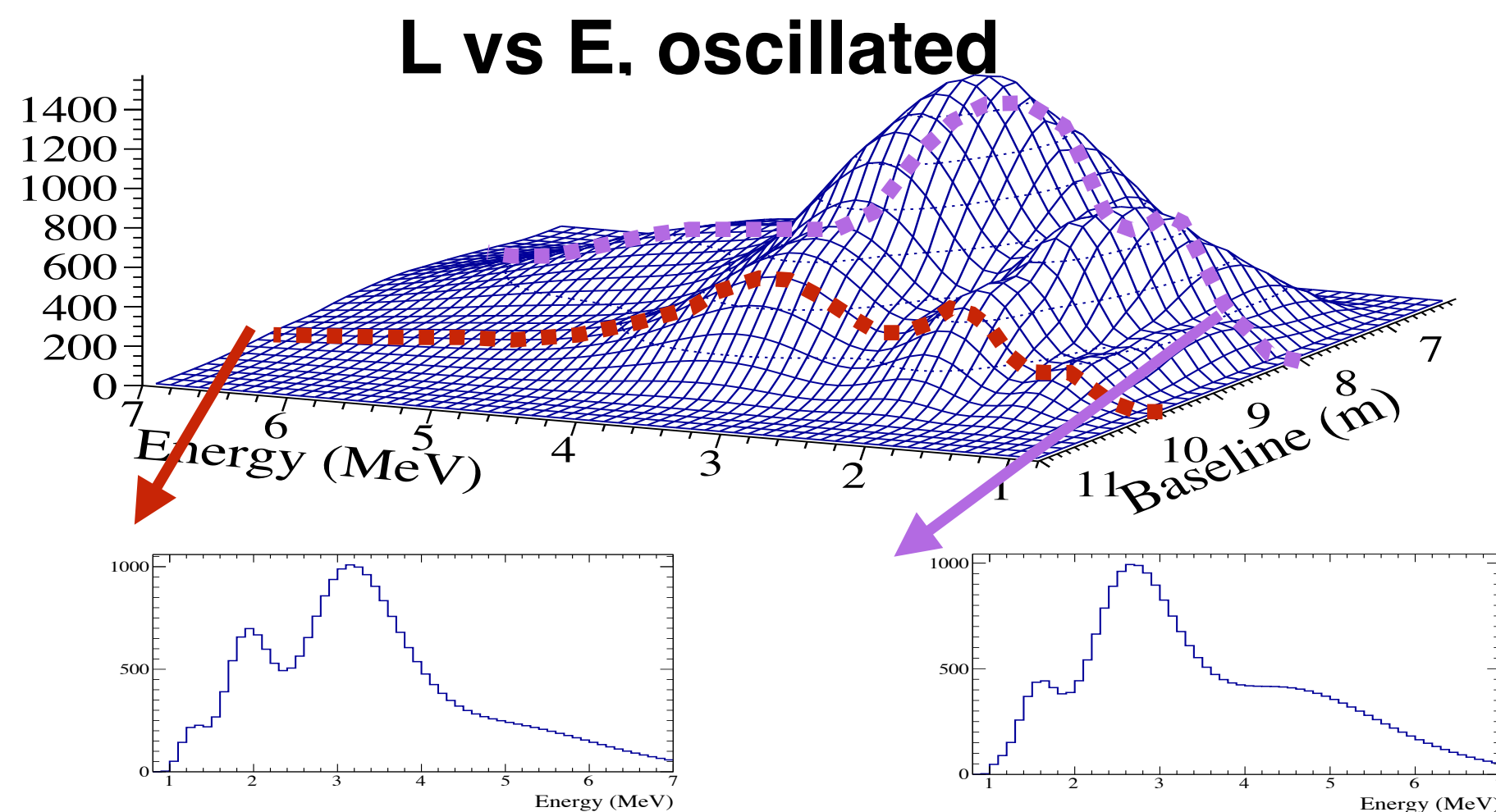
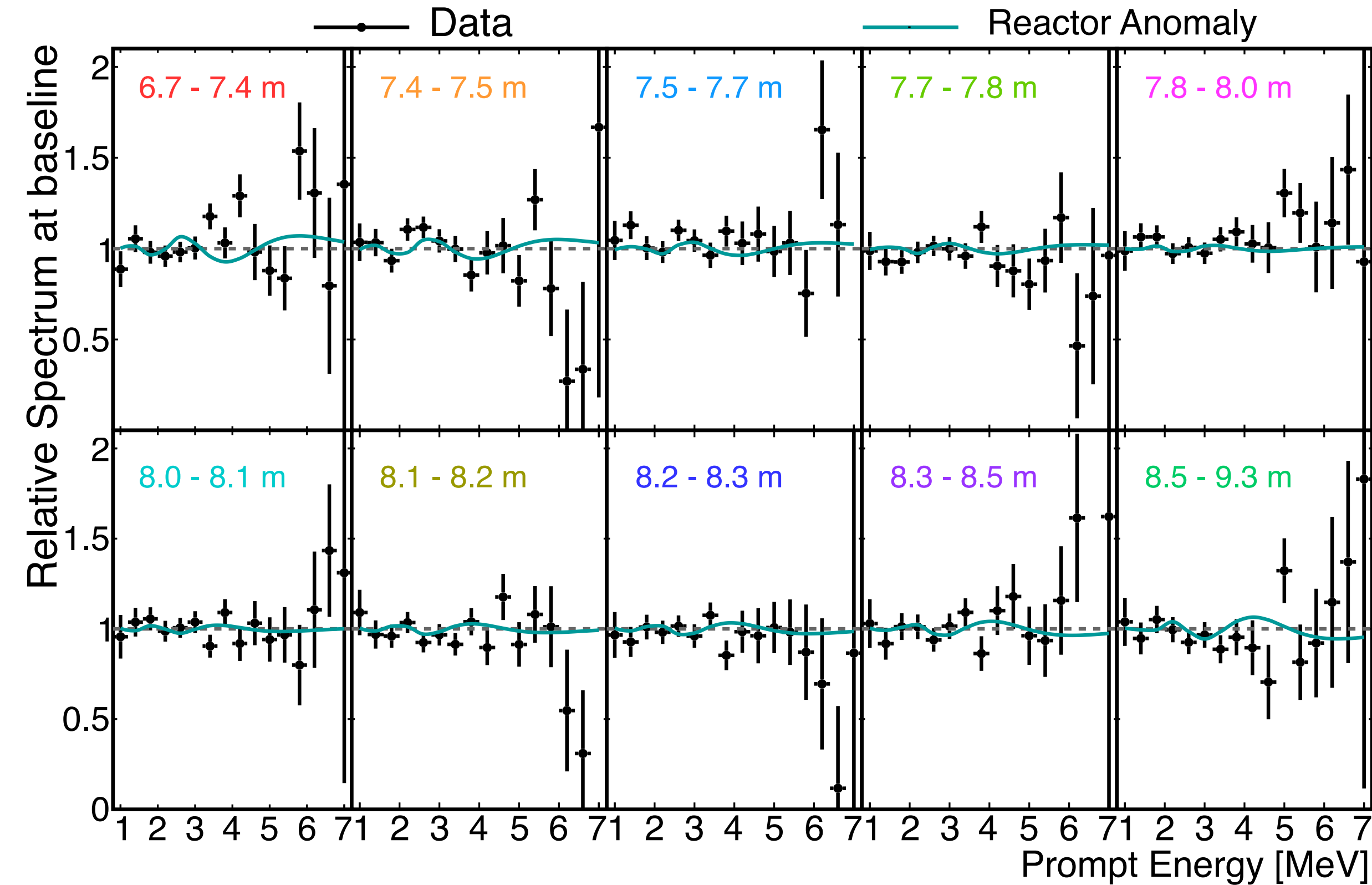
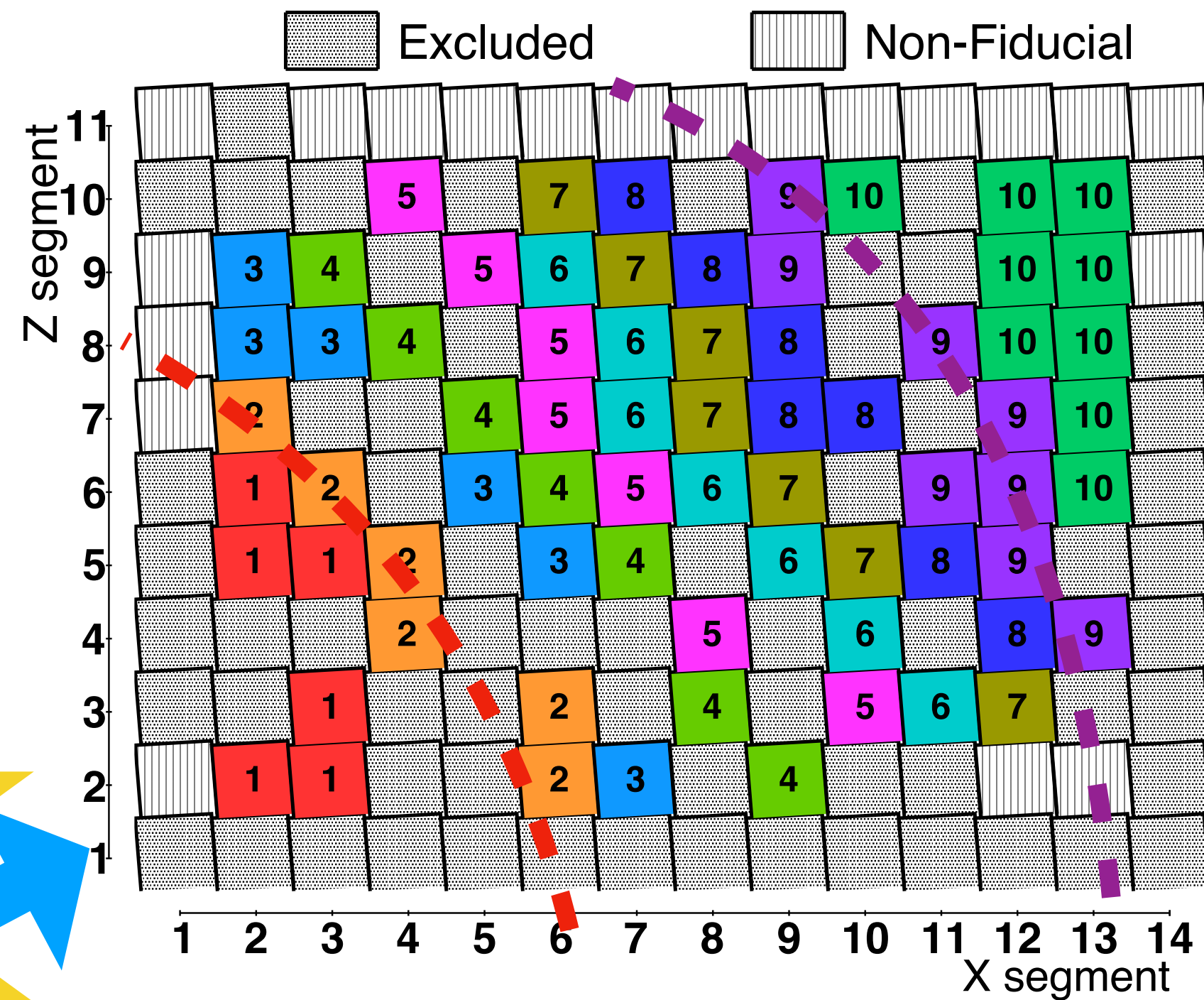
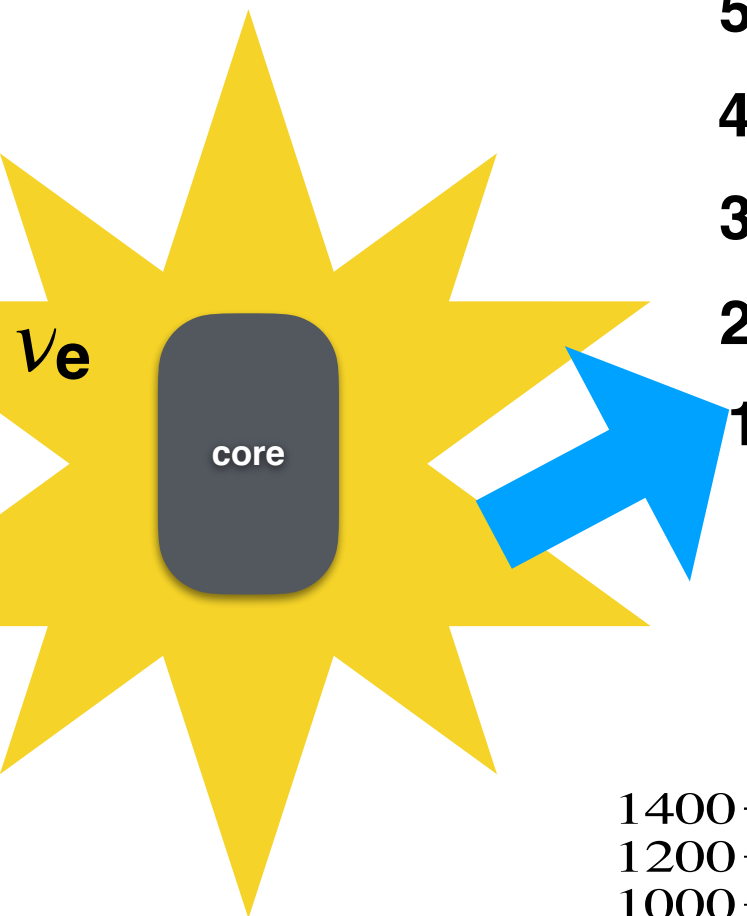


Reject candidates from 36 fiducial segments experiencing PMT current instabilities

- Time+position-coincident IBD e+ and n signals
- Prompt: IBD e+-like PSD+energy
- Delayed: n-<sup>6</sup>Li PSD+energy+topology
- Reject if coincident with cosmic  $\mu/n$
- Require signals to occur in fiducial segments
- Primary cosmic neutrons account for most of the remaining IBD-like background



# Oscillation Strategy



**No obvious deviations from flat no-oscillation scenario**

Oscillations modify energy spectrum as a function of baseline



# Oscillation Search: Results

- Compare measured, predicted spectrum ratios for different  $(\Delta m^2_{41}, \sin^2 2\theta_{14})$ :

$$\chi^2_{min}(\Delta m^2, \sin^2 2\theta) = \Delta^T \mathbf{V}_{tot}^{-1} \Delta$$

- Uncertainty covariance matrix  $\mathbf{V}_{tot} = \mathbf{V}_{sys} + \mathbf{V}_{stat}$

- Statistics are the dominant sensitivity limiter

- Best-fit  $\chi^2$ /NDF of 119.3/142 at  $(\Delta m^2_{41}, \sin^2 2\theta_{14}) = (1.78 \text{ eV}^2, 0.11)$

- Pictured:  $\Delta\chi^2$  with respect to this best-fit point

**Reactor Anomaly**

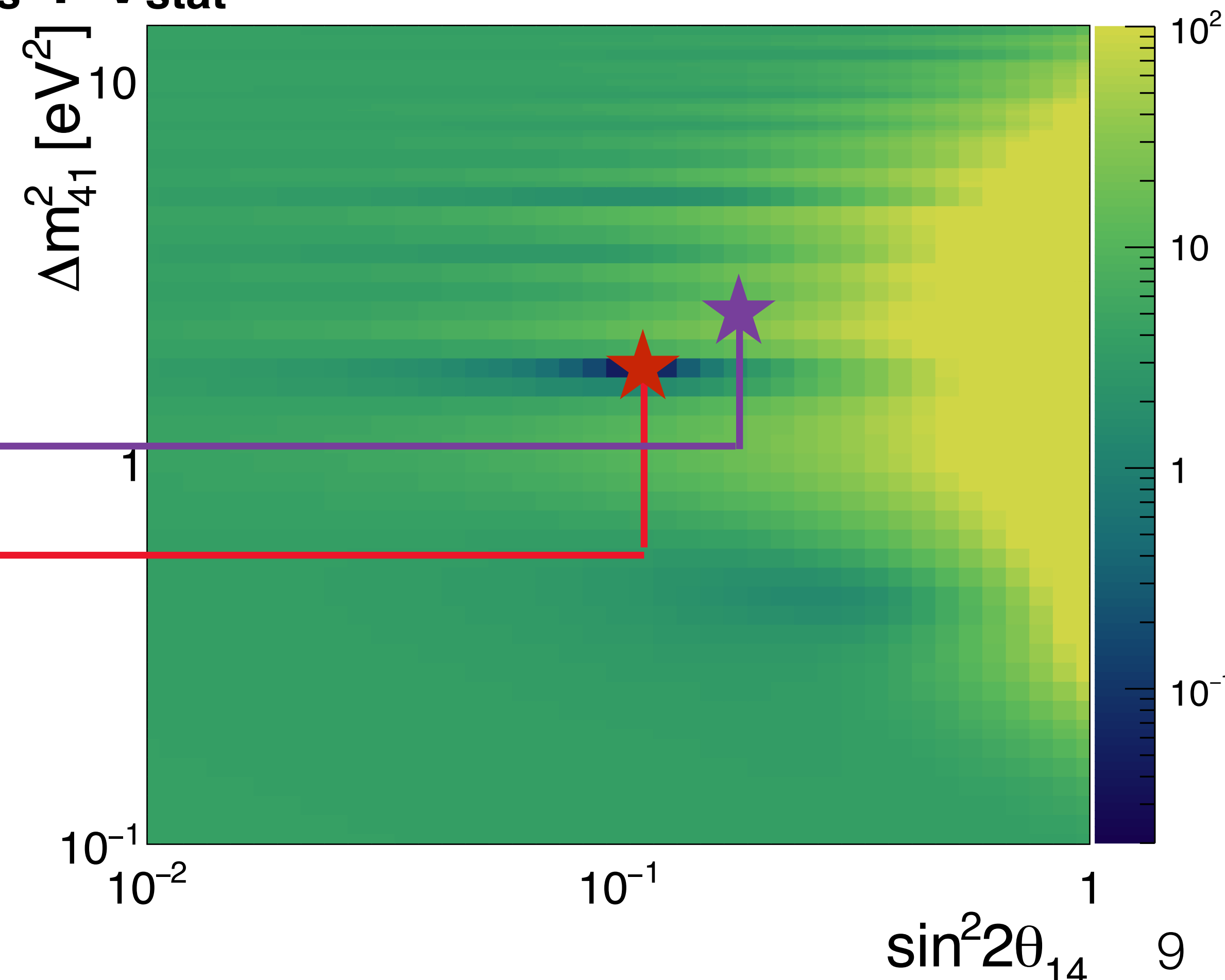
$\chi^2 = 135.1$

**Data**

$\chi^2 = 119.3$

**No oscillation**

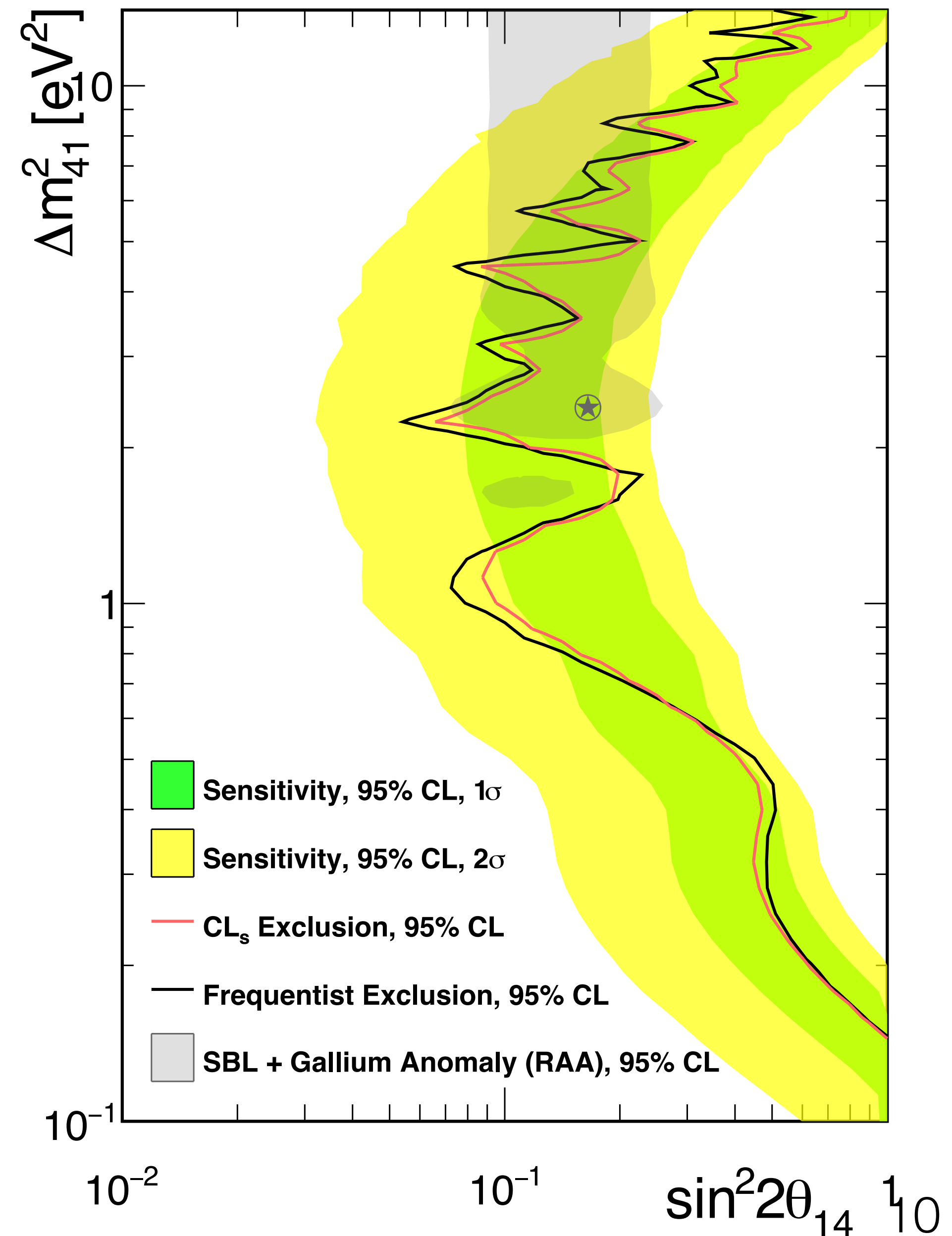
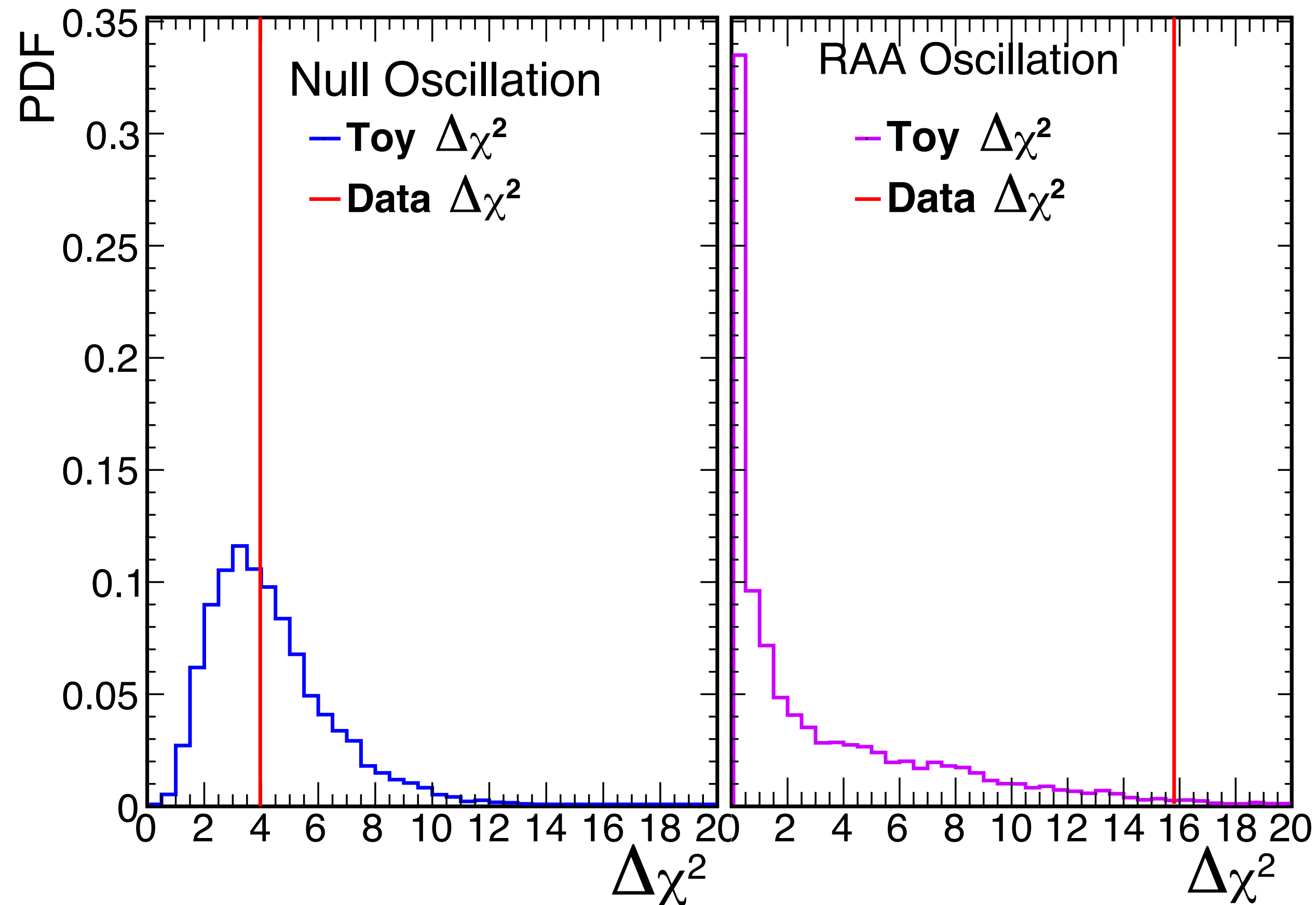
$\chi^2 = 123.3$





# Oscillation Search: Results

- Feldman Cousins frequentist test and Gaussian CLs method are used to evaluate the exclusion regions in the oscillation phase space.
- RAA best-fit excluded: 98.5% C.L.
- Data is compatible with null oscillation hypothesis ( $p=0.57$ )





# Summary

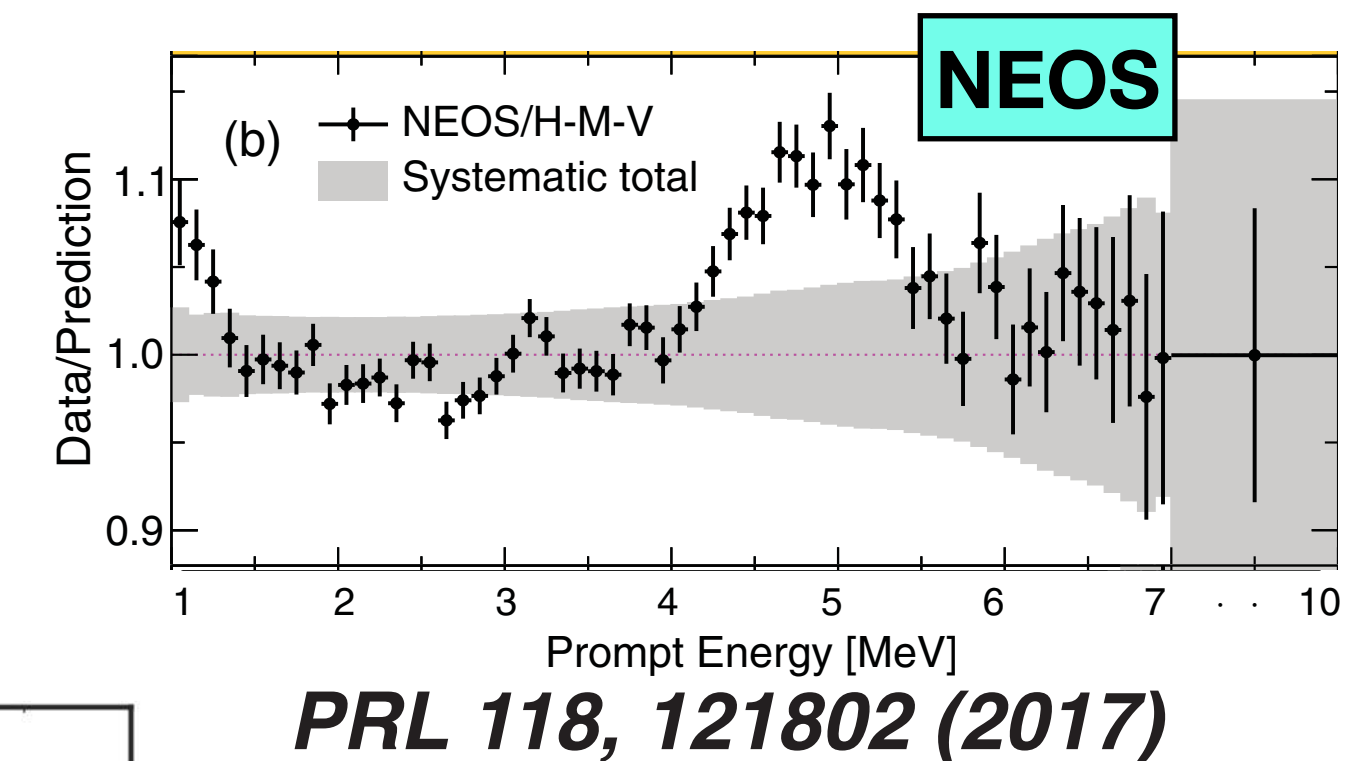
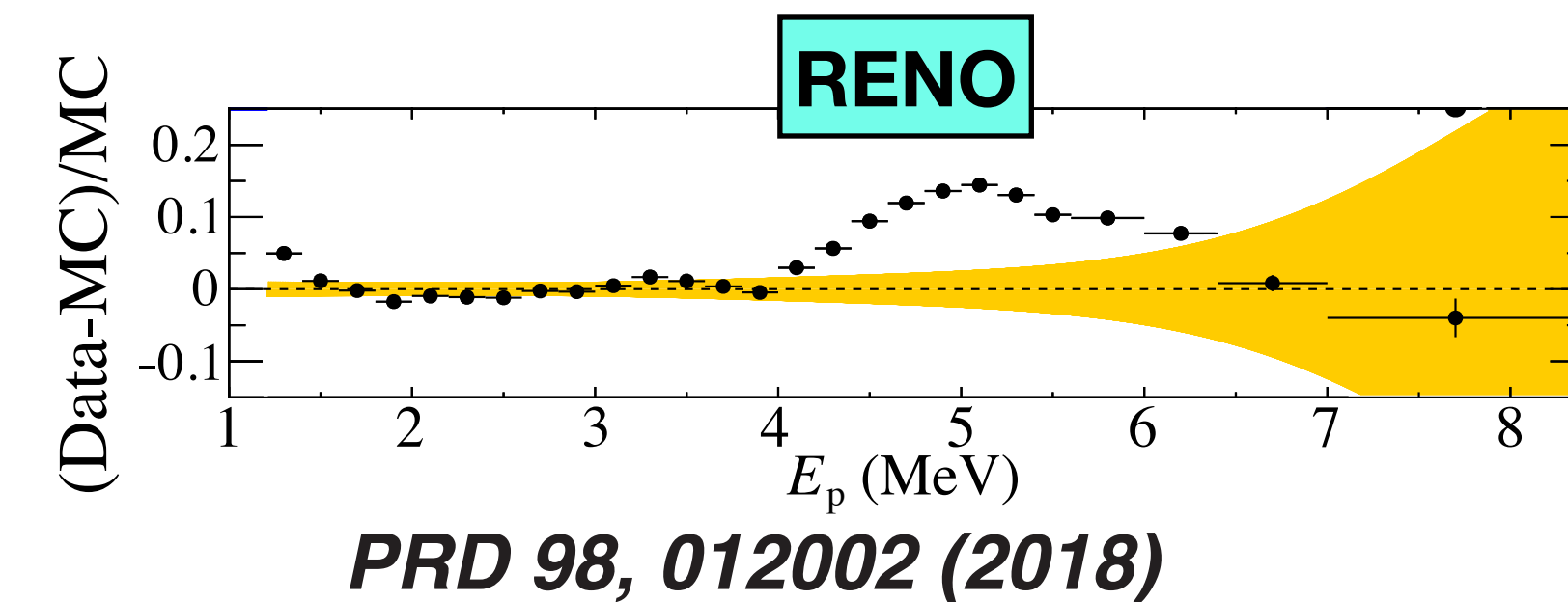
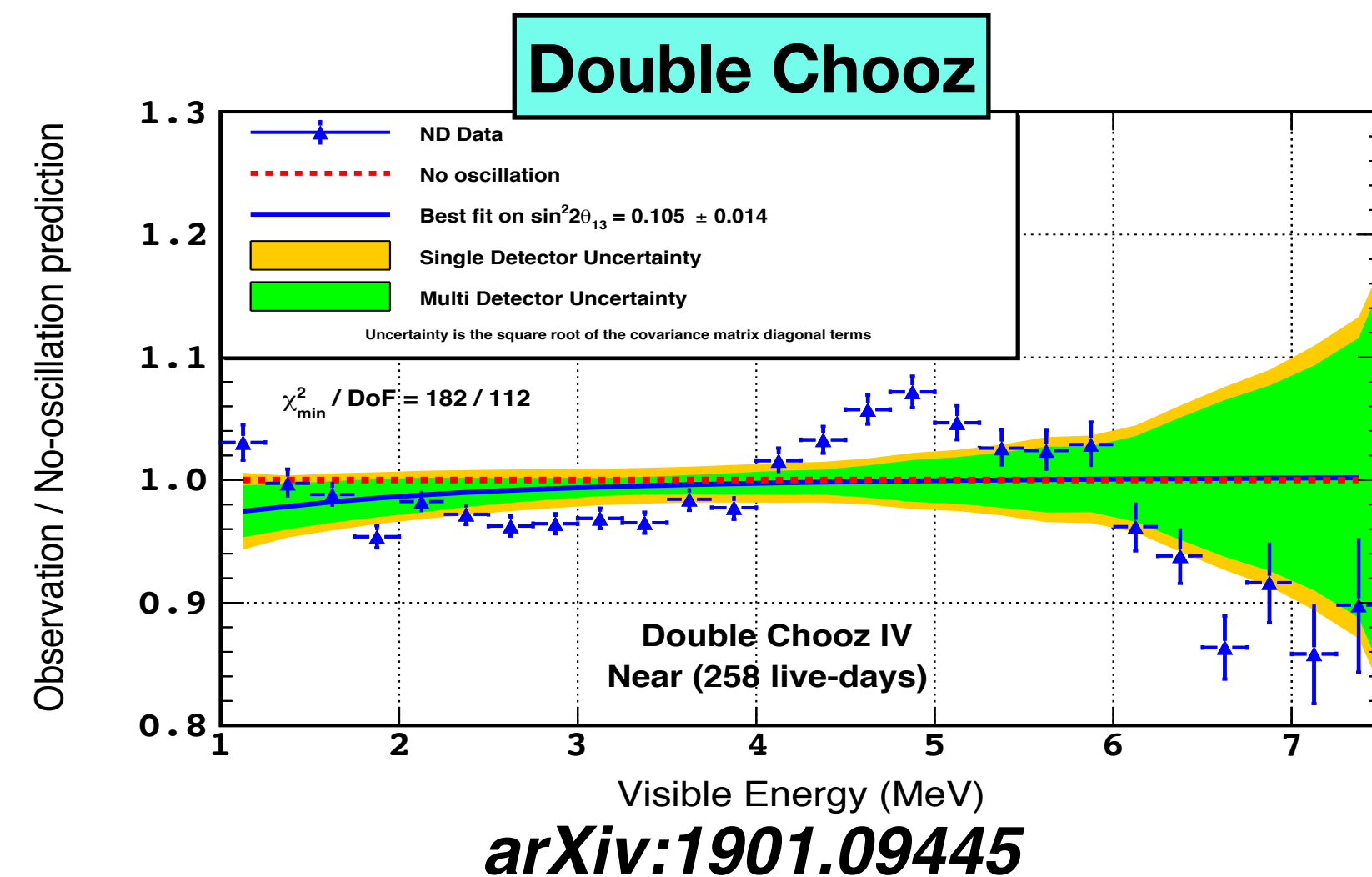
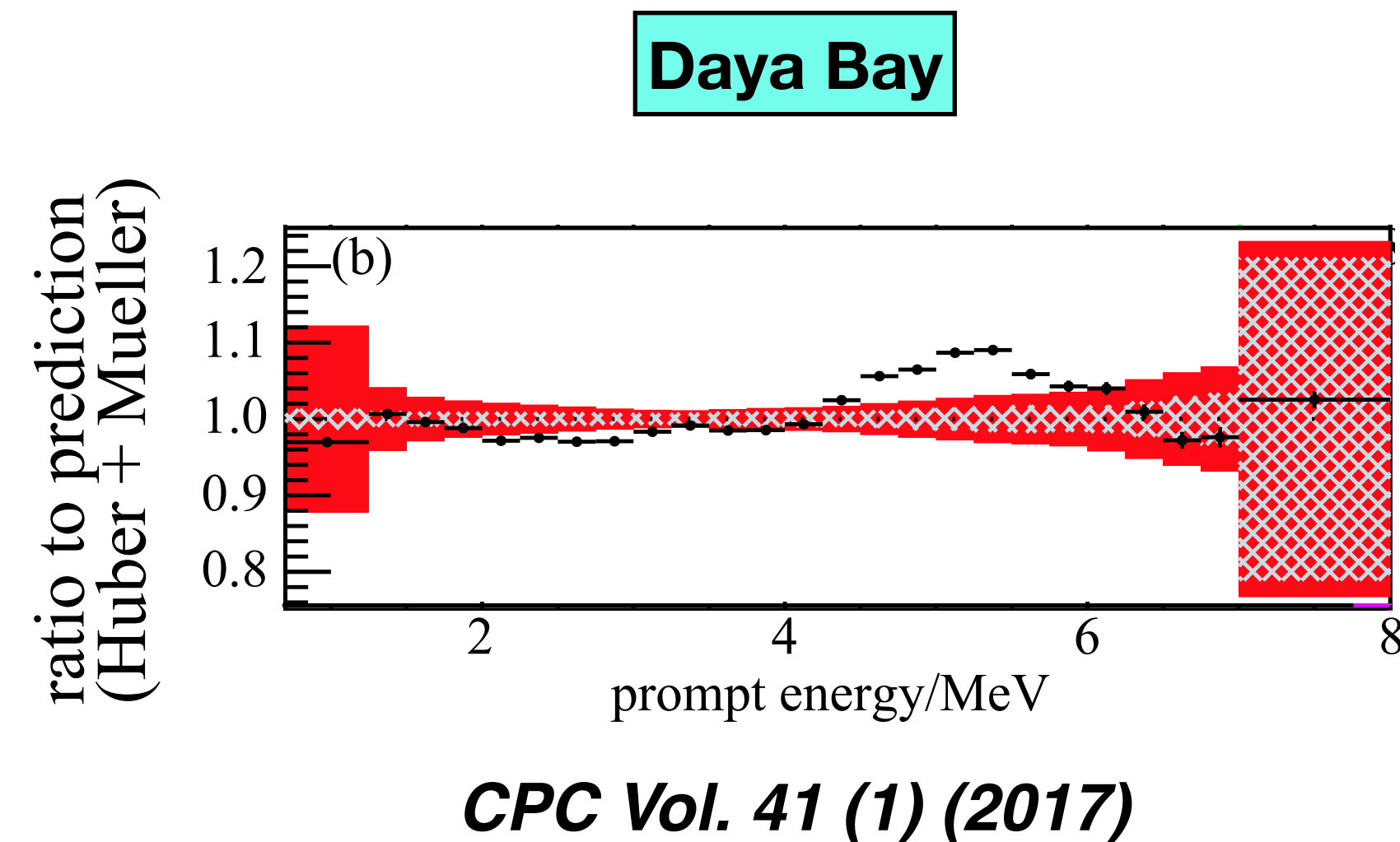
- An analysis of all PROSPECT reactor neutrino data has increased sterile neutrino sensitivity in the high- $\Delta m^2$  regime.
- The ‘reactor antineutrino anomaly’ best-fit is excluded at  $2.5\sigma$  CL.
- No evidence for sterile neutrino oscillations is found.
- PROSPECT’s current dataset will provide a substantially improved spectrum and oscillation measurement in the future. (EG.00004 X. Zhang)
- PROSPECT is pursuing upgraded detector deployment at HFIR that will further increase its measurement precision (EG.00005 H. Pieter)
- Latest Antineutrino Spectrum Measurement at PROSPECT (EG.00002 B. Foust)
- Check out these other PROSPECT talk(s) (EG00007 A. Delgado)



# Backup Slides

# Motivation: Reactor Antineutrino Spectrum Deviations

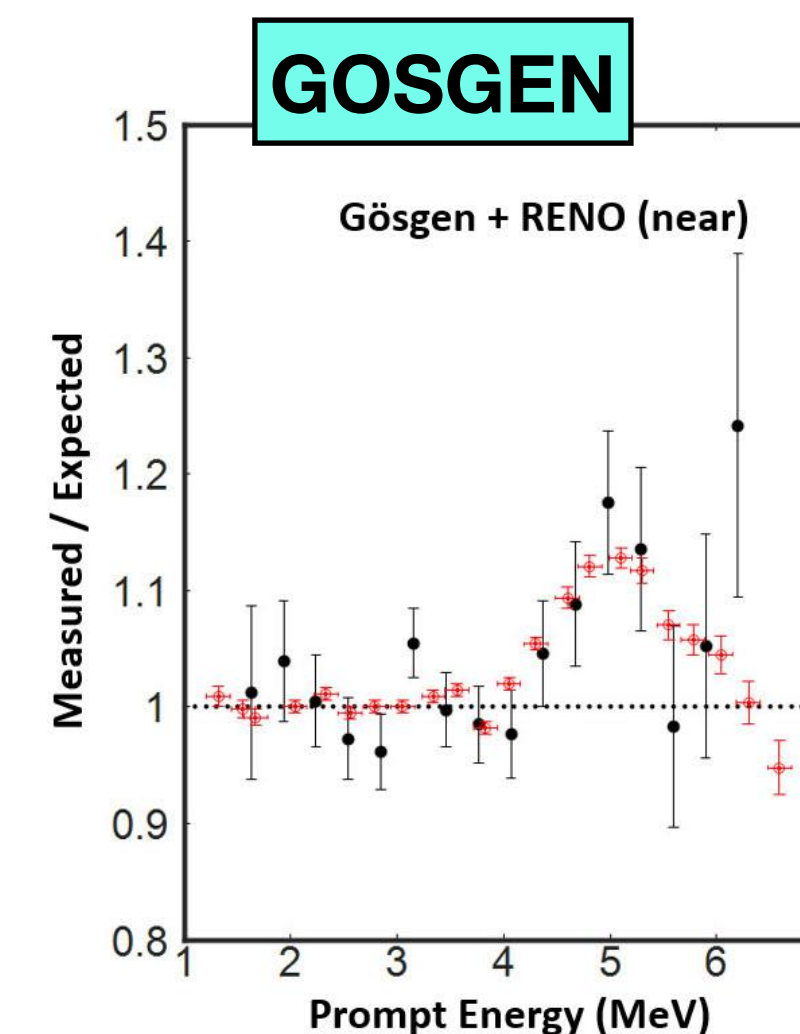
Experiments precisely measured spectrum from Low Enriched Uranium (LEU) reactors  $^{235}\text{U}$ ,  $^{238}\text{U}$ ,  $^{239}\text{Pu}$ ,  $^{241}\text{Pu}$



Distortion in 4-6 MeV prompt energy, not only on theta13 experiments.

## Where this deviation is coming from?

- Cannot be explained by the sterile neutrino introduced for flux deficit.
- Could be an issue with reactor models?  
Experiments used conventional reactors (LEU).



Re-evaluation (2018) of Gosgen(1980's) experiment also showed a deviation in 4-6 MeV region.

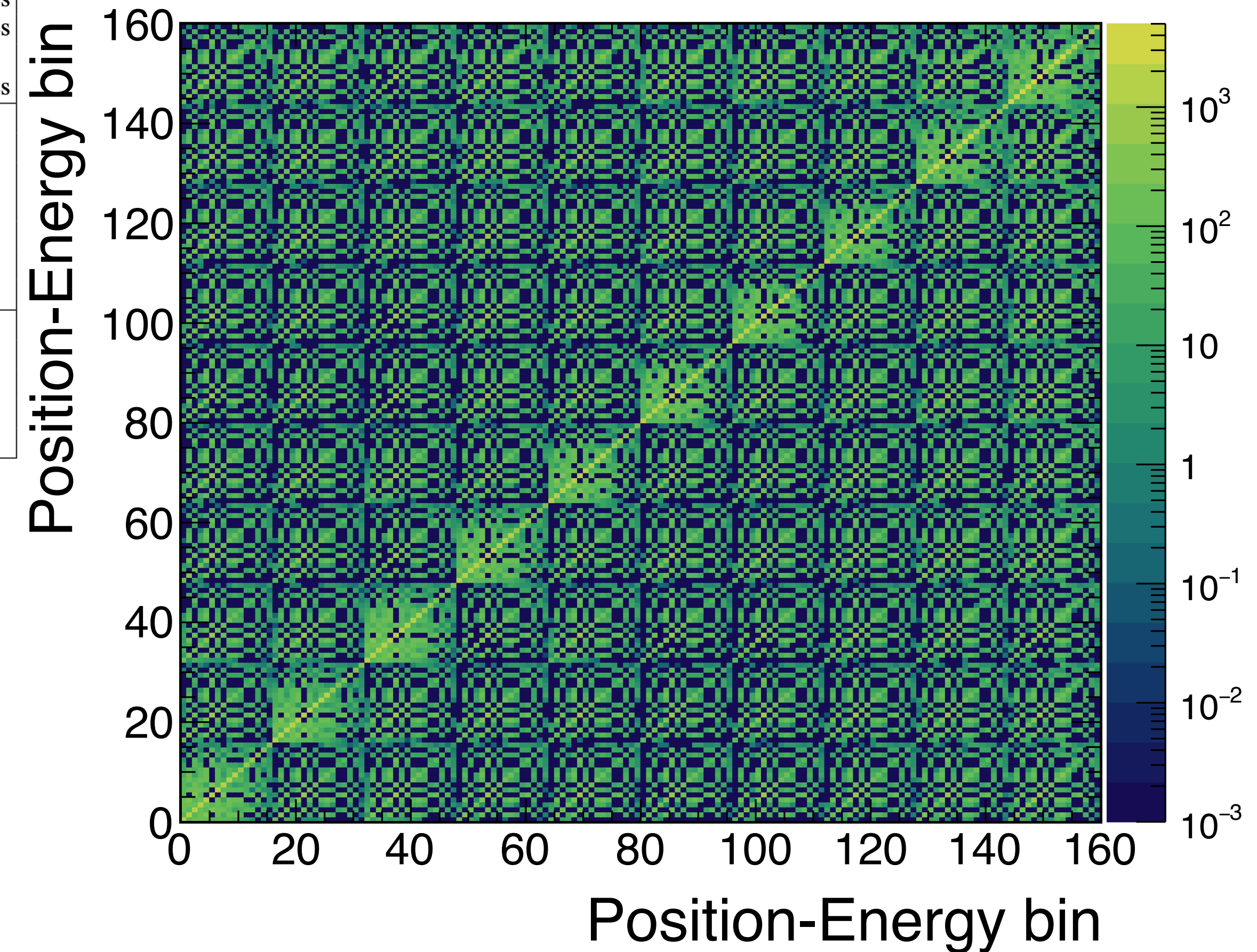
**arXiv:1807.01810**

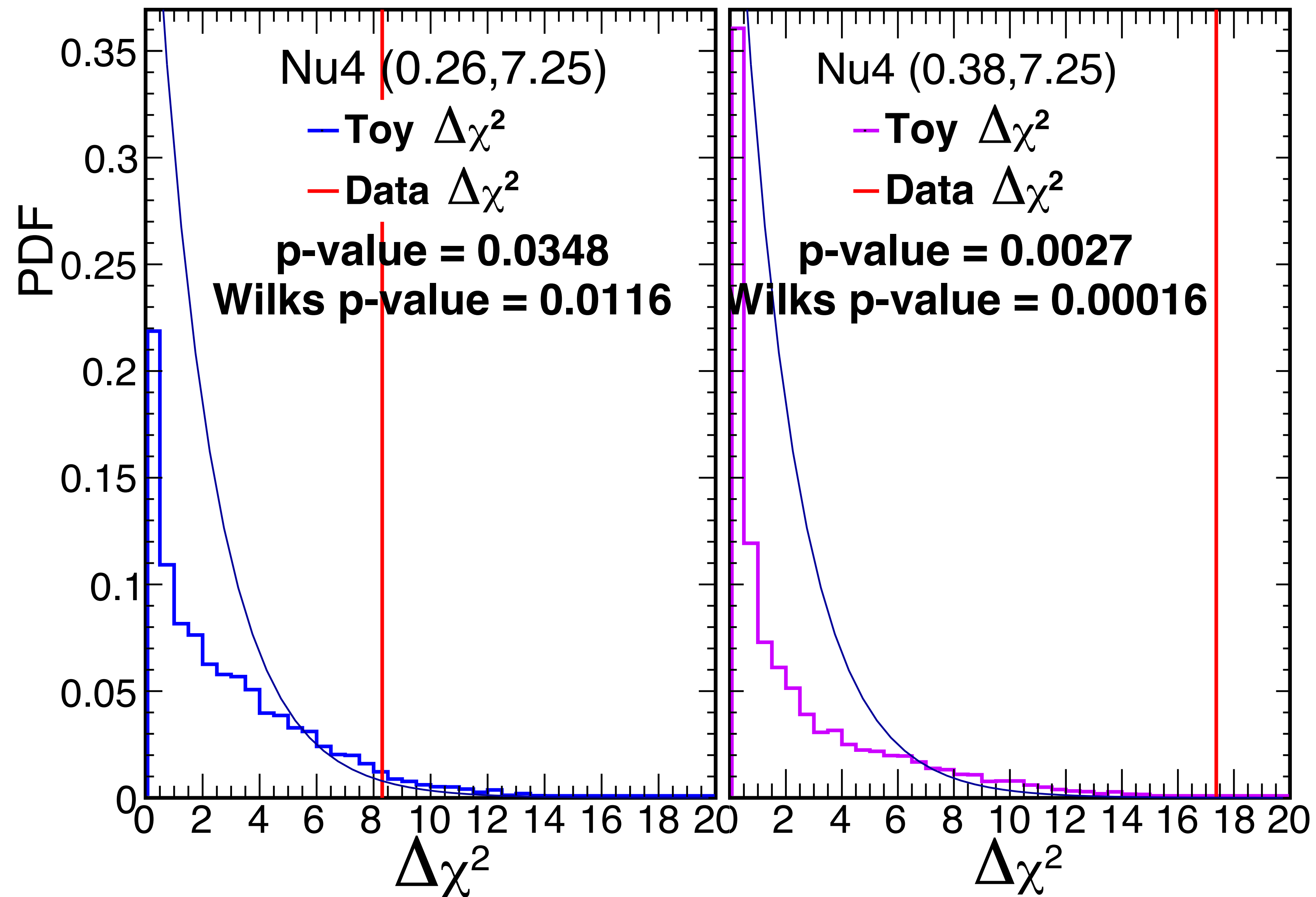


# Systematics

Parameter	Section	Nominal Value	Uncertainty	Correlations
Absolute background normalization	VI B, VI D	-	1.0%	Correlated between energies and baselines
Absolute $n$ -H peak normalization	VI D	-	3.0%	Correlated between energies and baselines
Relative signal normalization	V C	-	5%	Correlated between energies
Baseline uncertainty	II	-	10 cm	Correlated between energies and baselines
First-order Birks constant	IV B	0.132 MeV/cm	0.004 MeV/cm	Correlated between baselines
Second-order Birks constant	IV B	0.023 MeV/cm	0.004 MeV/cm	Correlated between baselines
Cherenkov contribution	IV B	37%	2%	Correlated between baselines
Absolute energy scale	IV B	-	0.6%	Correlated between baselines
Absolute photostatistics resolution	IV C	-	5%	Correlated between baselines
Absolute energy leakage	IV D	-	8 keV	Correlated between baselines
Absolute energy threshold	IV B, III G	-	5 keV	Correlated between baselines
Relative energy scale	III H, IV B	-	0.6%	Uncorrelated between baselines
Relative photostatistics resolution	III H, IV C	-	5%	Uncorrelated between baselines
Relative energy leakage	IV D	-	8 keV	Uncorrelated between baselines
Relative energy threshold	IV B, III G	-	5 keV	Uncorrelated between baselines
Reflector panel thickness	IV B	1.18 mm	0.03 mm	Uncorrelated between baselines

- The diagonal (statistical uncertainties) is clearly dominant
- Biggest systematics impact: relative segment normalization uncertainty, which effects low-dm2 values in particular







# Global

- PROSPECT and STEREO dominate  $> 3 \text{ eV}^2$
- DANSS and NEOS dominate at  $< 3 \text{ eV}^2$
- Full PROSPECT-II dataset will provide best coverage above  $\sim 1.5 \text{ eV}^2$

